That the exponent for scaling is about $1 / 4.1$ is suggested in an unexpected area. The distribution of meteorite falls by size and number was found by H . Brown ${ }^{28}$ to follow a relation $\Delta \mathrm{log}$ mass $/ \Delta \log$ frequency $=0.77$ through the range of 32 kg to 1,024 kg . He also pointed out the "striking resemblance" of the frequency distribution curve as a function of magnitude for asteroids (see Kuiper, et al., Ref. 19) to the curves for meteorites, the congruent slope being 0.76 . A study by Opik $^{20}$ of the size distribution of 812 craters on Mare Imbrium when plotted in the same manner, except that a fixed 20/1 ratio of crater to meteor diameters is used, gives a slope between $0.45-0.50$. However, applying the energy exponent $1 / 4$ instead, the slope is $0.72-0.75$. All this may be fortuitous, but it does represent different kinds of sampling of large numbers of the same kind of things in the space around the earth.

The frequency, in fact the very existence, of random bodies in the $300-600 \mathrm{~km}$ range may well be zero, and in the present inventory of astronomers, the number is zero. However, the presence of large maria on the moon is disturbing in this regard. It may be that in our time lunar explorations will clear up the dating and mechanism of formation of maria, but until then it should be useful as well as instructive at least to consider the possible effects of large collisions.
Penultimate disasters have not been missed by far in time or distance. The "mythologies" of ancient peoples include descriptions of very close passes. In 1960 an unknown object of an estimated 8 cubic km passed close to our orbit, 8 million km from the earth. The asteroid Hermes slid to within $320,000-640,000 \mathrm{~km}$ of the earth in 1937, and near the turn of the century, Eros, later Amor, came within 20 and 16 million km . Estimated kinetic energies represented by these bodies are in the order of $10^{30}, 10^{20}, 10^{31}$ and $10^{32}$ ergs. Visitations by these would have been no myths.

## ACKNOWLEDGMENTS

I am indebted to A. O. Kelly, friend and collaborator, for the sustained interest in this field through years of correspondence. Colleagues and associates have contributed through hard criticism and ideas. Dr. P. D. Krynine and Dr. V. Vand have been especially helpful. Any errors are my own.

## REFERENCES

1. L Coes, Jr. A new dense erystalline silien, Science, 118, 131, 1953.
2. F. Dachille and R. Koy. High pressure region of the silica isotypes, Z. Krist., 111, 451-61, 1959
3. E. C. T. Chao, E. M. Shoemaker and B. M. Madsden. First natural occurrence of coesite, Science, 132, 220-222, 1900.
4. E. C. T. Chno, J, J. Fahey, J Littler and E. J. Milton. Stishovite, SiO $\mathrm{O}_{2}$, a high pressure mineral from Meteor Crater, Arizona, J. Geophys, Res., 67, 419, 1962.

4a. S. M. Stish 10, 837-39, 1 !
4b. J. J. Fahe personal cor
5. H. H. Nini 1956.

5а. H. H. Ninf Western Au
6. R. S. Dietz
7. V. B, Meen

8, C, S. Benls. Fergubon a
9. A. O. Kell: Earth Scie College, Per
10, G. H, Wn Rieacя, Doc
11. T. Kaljuve Tallin-Reva
12. V. Vand. I
13. C. S. Richt 1958.
14. H. Shapley
15. Cratering : liminary ty Research, 6
16. G. Fielder. structure o 123, 15-26,
17. R. B. Vail 1961.
18. H. Brown. borhood of
19. G. P. Kull Ser.), 3, 2 !
20. E. J. Opil 34-36, 1957.
21. D. M. Bar 30, 23-28, 1!
22. A. J. Ear Now York.
23. J. D. Boor tonship to
24, J, D, Boon Laboratory

GENERALR. B. Bald B. Gutenbe

